

Quantum Well Infrared Photodetectors for Astronomy

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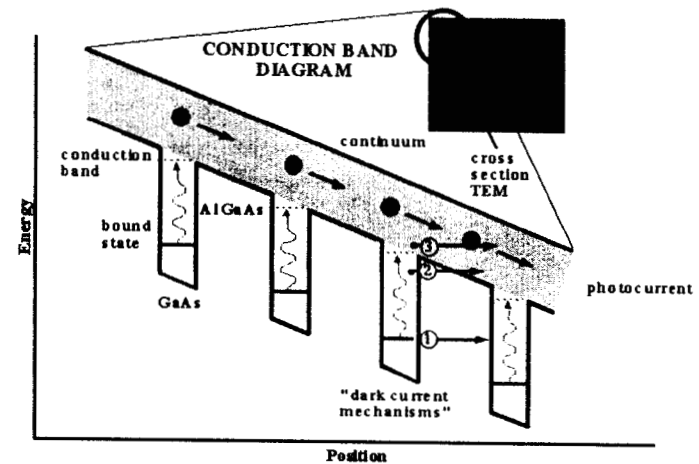


QWIP Properties

Realization of quantum wells in AlGaAs

Advantages:

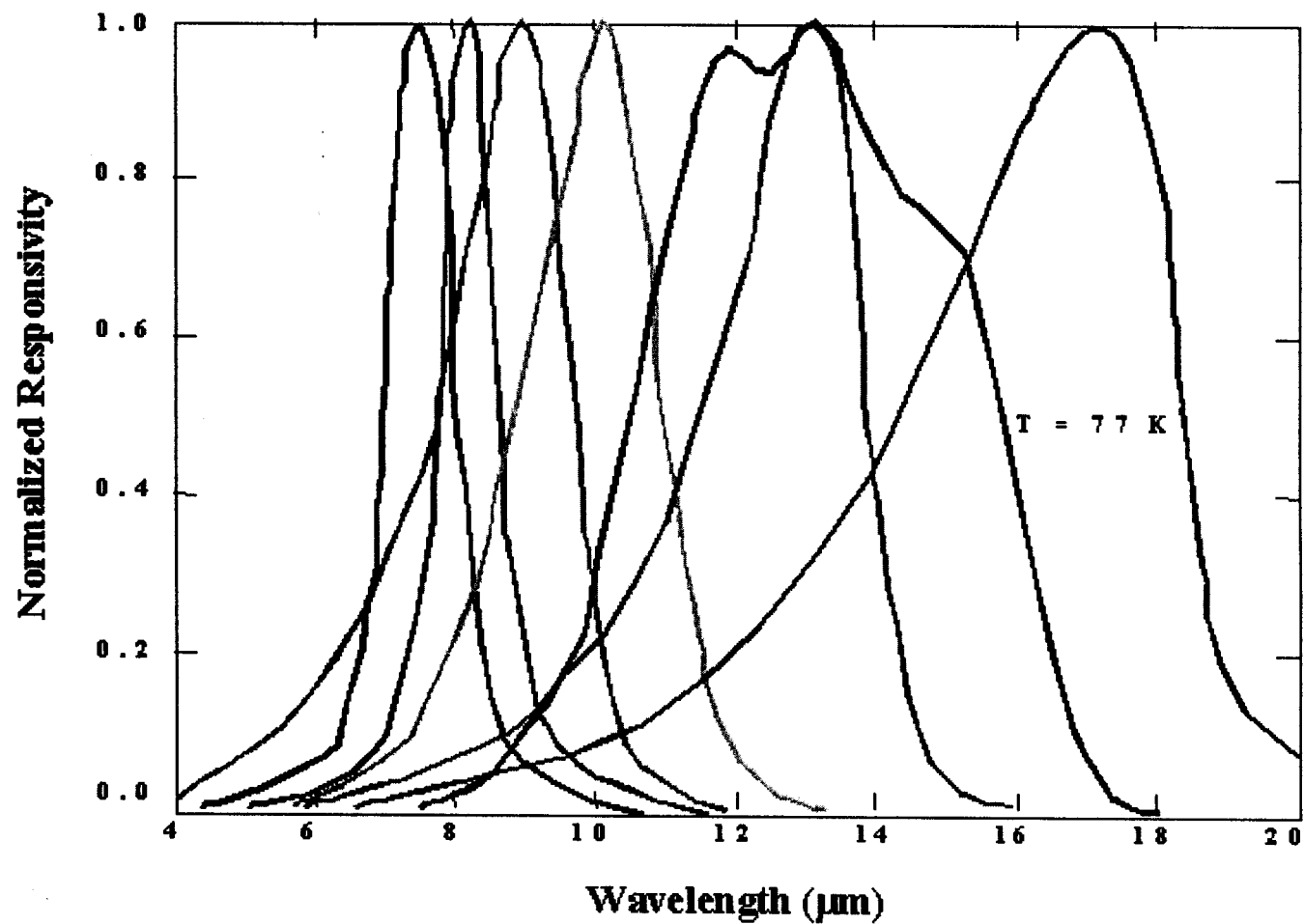
- Tunable bandpass
 - 15 μm device for surveillance applications
 - 8.5 and 12.5 μm devices in QWICPIC, astronomical camera at Palomar
 - devices could be doped for far-IR wavelengths $\geq 70 \mu\text{m}$
- Higher temperature operation
 - 25–30 K for low background applications
- Low 1/f noise
 - None measured to 30 mHz
- Easily mated to existing CMOS muxes
 - 640 \times 480 already produced
 - 1024² in development



Disadvantages

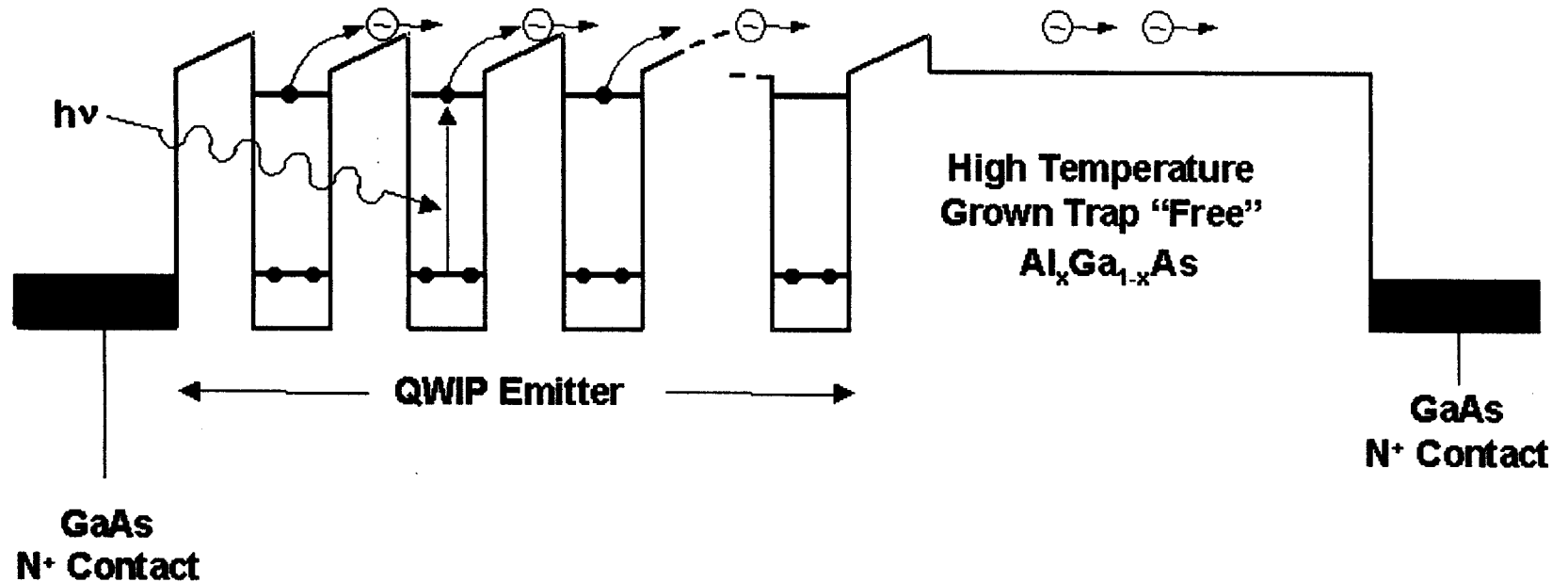
- Narrowish bandwidth
 - 10% for typical structure
 - “broadband” devices still only 50%
- Modest quantum efficiency
 - Photons must enter \parallel to surface
 - Diffusers mated to detector improve performance
 - Best QE's are of order 25%

Spectral Response



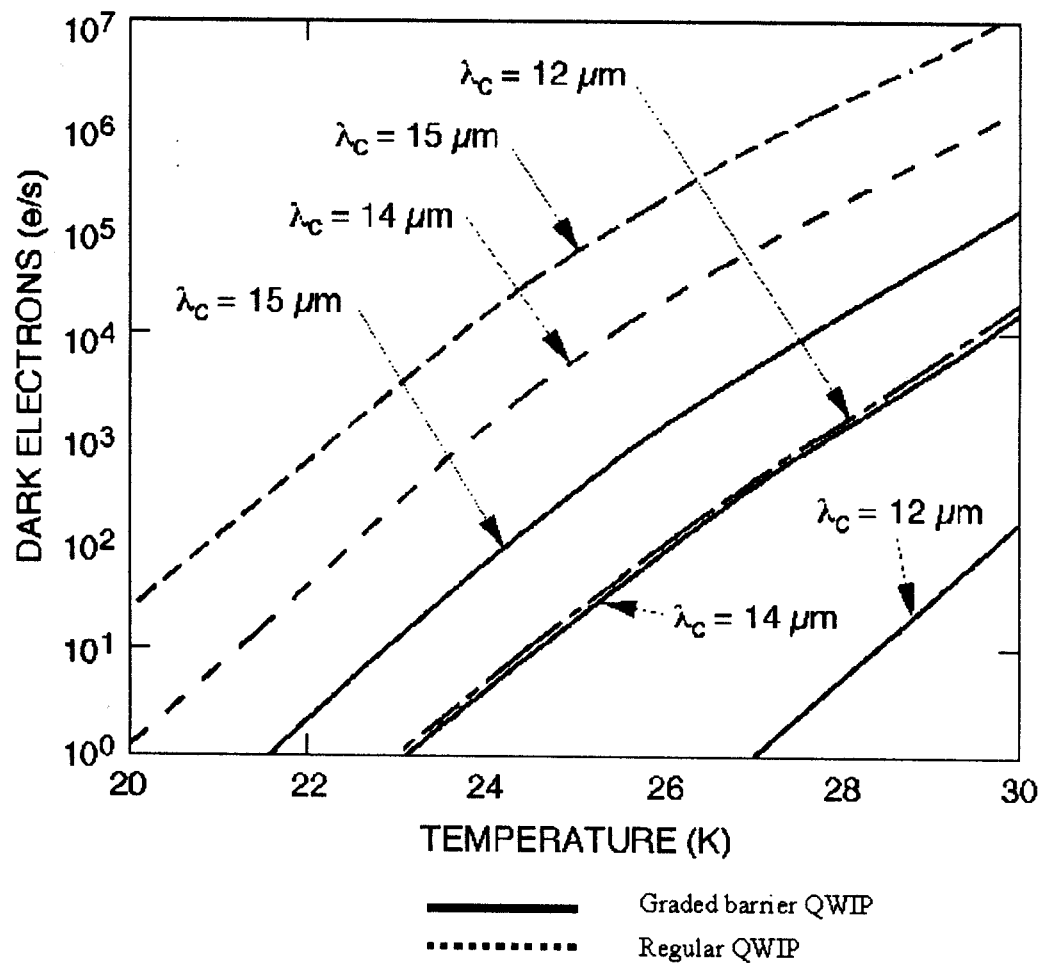
Spectral response of a suite of classical QWIPs along with a broadband version.

Overcoming Low Flux Problems



- Simple QWIPs suffer same electric field problems as photoconductors at very low backgrounds.
- Solution: BIB-like structure, with trap-free blocking layer

Dark Current Properties



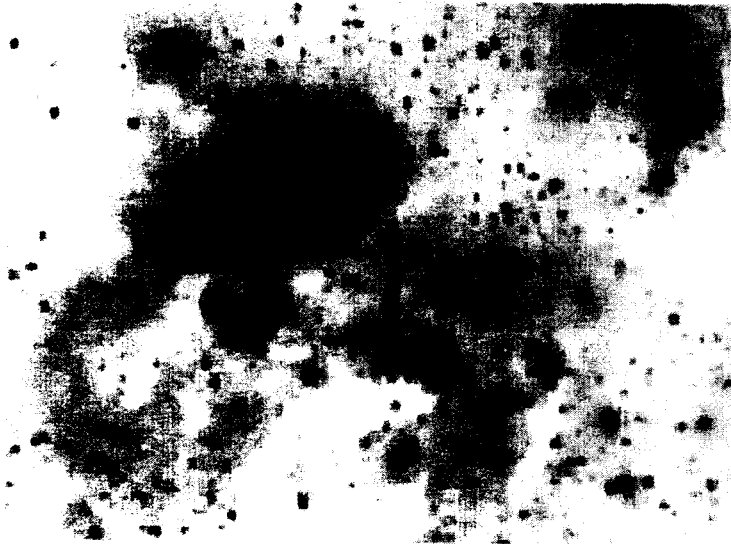
Estimated dark current properties of a $40 \mu\text{m}$ square pixel.

Orion Trapezium

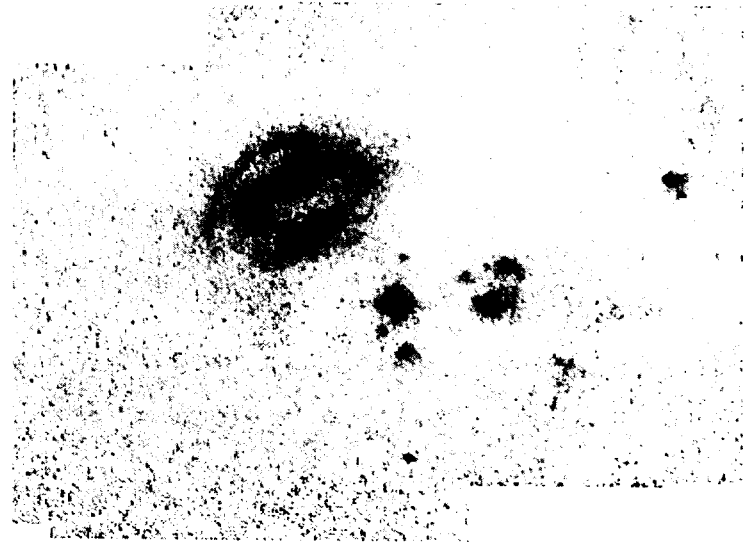


- QWICPIC has been at Palomar for 2 years.
- 256^2 $8.5\ \mu\text{m}$ QWIP
- Observations of high mass star forming regions
- Will soon install $12.5\ \mu\text{m}$ 640×480 QWIP for dual-band imaging

W3 Star Forming Region



2 μm



8.5 μm

QWICPIC has provided detailed information about temperature structure in ionized areas in star forming region.

Summary

- QWIPs are a viable alternative when 10 K cooling is not available
- Tunable passbands can be optimized to mission
- Ultra low $1/f$ noise allows long, stable integrations
- Easily manufactured and can be scaled to large formats
- In production now, and working in an astronomical environment